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Thermochemical modelling of the subsurface environment on Enceladus

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The subsurface environment of Enceladus is potentially habitable: there is a global subsurface ocean [1], energy from hydrothermal activity [2] and bioessential elements [3]. Carbon, as a fundamental bioessential element, it is critical for life, so understanding how it is processed within the Enceladus environment is crucial in assessing Enceladus’ potential habitability. Carbon is likely to be bound within the silicate interior [4] and liberated through water-rock (silicate-ocean) interactions.

We have undertaken thermochemical modelling (CHIM-XPT) [5] of these interactions and tested different hypotheses for the formation of Enceladus. Both models reacted the silicate interior (with a CI chondrite composition [6]) with a fluid representative of the subsurface ocean: a) a dilute sodium chloride solution, based upon the assumption that the subsurface ocean originated as almost pure water [7]; b) a solution with a cometary composition based upon data collected from 67P [8], based upon the assumption that the water originated from melted cometary ice [9]. We have explored the full temperature and pressure ranges anticipated at the rock-water interface [2].

We will present the outcomes from this modelling, which includes a theoretical composition for a modern day subsurface ocean, potential carbon cycling pathways and the effect of carbon species on the pH of the subsurface ocean fluid.