Extra-solar planetary atmospheres and interiors

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Extra-solar Planetary Atmospheres and Interiors

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Summary:
This project aims to bring together models of both the atmospheric circulation and interior structure and apply them to close-in gas giant exoplanets, with the intent of producing information that can be used to inform and interpret observations. To this end, an atmospheric model has been modified, and an interior model is being adapted. Scans through parameter space are being performed with the atmospheric model to validate the modifications.

Models:
The atmospheric model used is PUMA¹, maintained at the University of Hamburg and based on the Reading spectral model. The atmosphere is modelled as a thin spherical shell under the hydrostatic approximation, and the hydrostatic primitive equations are solved.

Moments of PUMA have been modified in this project to allow other planets to be studied. An interior model is also under development, scaling down a model for low-mass stars and brown dwarfs. This is based on an expanded and upgraded version of the code by Mazzitelli².

Conclusions:
PUMA has carried out preliminary scans through parameter space. As the rotation rate, \( \Omega \), decreases, the jets and Hadley cells expand towards the poles, demonstrating the predictions of the Held-Hou model³, which gives the poleward extent of the Hadley cells as proportional to \( 1/\Omega \). Super-rotation also becomes apparent.

For a smaller planet, the jet velocity is seen to increase, again as predicted, as the wind at higher latitudes is given as proportional to \( 1/a \), where \( a \) is the radius of the planet.

Outlook:
The next stage will involve focusing the model toward the specific characteristics of a small number of chosen exoplanets. In parallel, the development of an interior model for these planets will continue. Although the models cannot be directly coupled, a link will be formed via the boundary conditions, using the output of each model to inform input to the other.

References: